

PHY 564

Advanced Accelerator Physics

Lectures 20

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Lecture 20. Full 3D matrices. Chromatic effects. Longitudinal (energy and time, synchrotron) oscillations in storage rings.

While we considered in many details 2D (transverse) matrices, we left aside a more complicated (and heavier) full 3D linear matrices and stability in periodic systems. Here is a brief recollection of what can be done (usually by computers) for linearized motion in arbitrary accelerating system.

First, let's remember equation of motion for the reference particle, e.g. that whose trajectory

$$\begin{aligned}
 \vec{r}(s) &= \vec{r}_o(s); \quad f' \equiv \frac{df}{ds}; \quad \vec{\tau} = \vec{r}_o'(s); \quad \vec{\tau}' = -|\vec{\tau}'|\vec{n}; \quad \vec{b} = [\vec{n} \times \vec{\tau}] \\
 \vec{r} &= \vec{r}_o(s) + x \cdot \vec{n}(s) + y \cdot \vec{b}(s); \\
 \frac{d\vec{\tau}}{ds} &= -K(s) \cdot \vec{n}; \quad \frac{d\vec{n}}{ds} = K(s) \cdot \vec{\tau} - \kappa(s) \cdot \vec{b}; \quad \frac{d\vec{b}}{ds} = \kappa(s) \cdot \vec{n}; \\
 P_1 &= P_x; \quad P_2 = (1 + Kx)P_s + \kappa(P_x y - P_y x); \quad P_3 = P_y;
 \end{aligned} \tag{20-1}$$

we use as a reference ($x=y=0$, $P_x=P_y=0$), whose arrival schedule

$$t(s) = t_o(s) \tag{20-2}$$

and energy (momentum)

$$cp_o(s) = \sqrt{E_o(s)^2 - m^2 c^4} \tag{20-3}$$

we follow.

We should note that the Frenet-Serret coordinate system is uniquely defined when curvature of the reference trajectory:

$$K(s) \equiv \frac{1}{\rho(s)} = \left| \frac{d^2 \vec{r}_o}{ds^2} \right| \equiv |\vec{\tau}'|. \quad (20-4)$$

is non zero. It also defines the direction of the normal and by-normal vectors \vec{n}, \vec{b} :

$$\vec{n} = -\frac{\vec{r}_o''}{|\vec{r}_o''|}; \vec{b} = \frac{1}{|\vec{r}_o''|} [\vec{r}_o' \times \vec{r}_o'']; \quad (20-5)$$

Note that curvature in (20-4) is positively defined. It means that even for a wiggly planar trajectories one either should use torsion (local rotation) or alternating sign of the curvature – we are using the later.

When curvature is zero (a straight line piece of the reference trajectory), both \vec{n}, \vec{b} are not uniquely defined, e.g. we can use torsion (rotation in \vec{n}, \vec{b} plane) as an instrument. We already used it for calculating matrices of solenoid and SQ-quadrupole. The only one important condition remains: you start from fully defined the \vec{n}, \vec{b} at the end the curved section, turn them around as many times and in any direction your want, but that you put the vectors \vec{n}, \vec{b} into the required directions at beginning of next curved section.

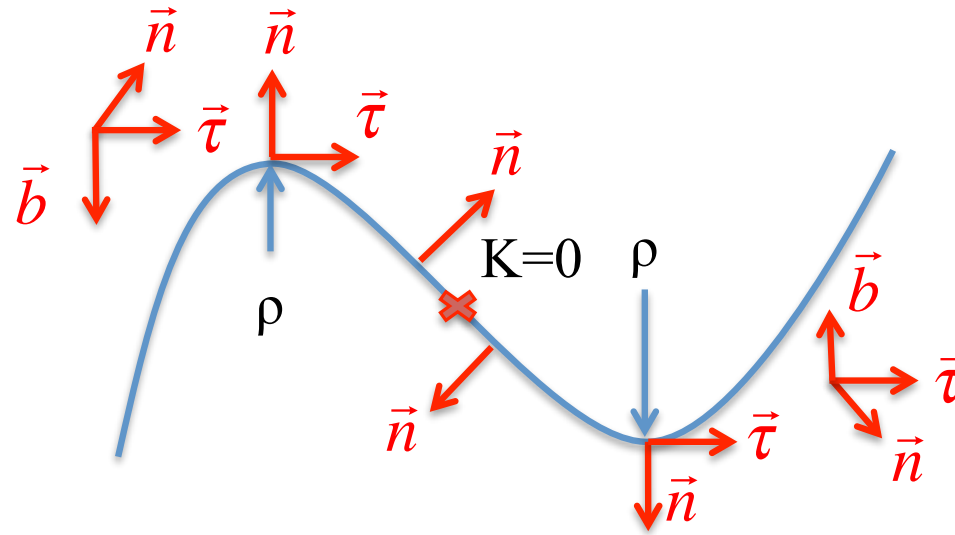


Fig.20-1. Even for a plane trajectory, the curvature direction \vec{r}_o'' can change, at some point (or at the straight line) the direction of the normal and bi-normal vectors has to flip (180-degrees rotation about $\vec{\tau}$) Hence, for such reference trajectories there must be either non-zero torsion (rotation) or provision for alternating sign of curvature. For planar trajectory defining direction of the bi-normal vector is sufficient to determine the sign of the curvature – usually it is selected to aim from the plain to the viewer.

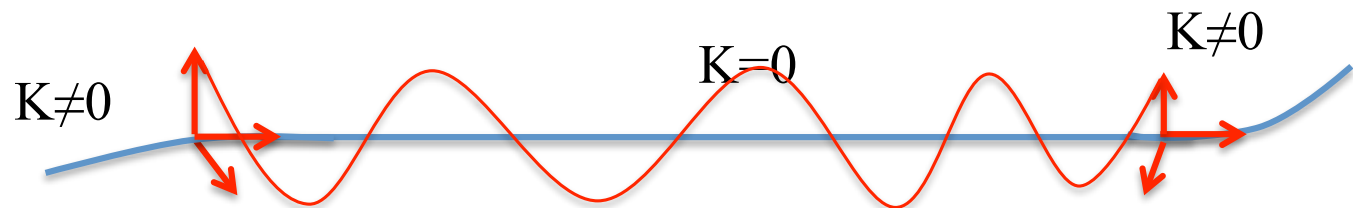


Fig. 20-2. Allowed manipulations with \vec{n}, \vec{b} in a straight section: it can be any helix with variable thread, but with two unit vectors \vec{n}, \vec{b} perpendicular to each other $\vec{b} = [\vec{n} \times \vec{\tau}]$.

We derived the conditions for the reference particle with

$$\vec{r} = \vec{r}_o(s); \quad t = t_o(s); \quad H = H_o(s) = E_o(s) + \varphi_o(s, t_o(s)) \quad (20-6)$$

as:

$$\begin{aligned} K(s) &= -\frac{e}{p_o(s)c} \left(B_y(0,0,s,t_o(s)) + \frac{c}{v_o} E_x(0,0,s,t_o(s)) \right); \\ B_x(0,0,s,t_o(s)) &= \frac{c}{v_o} E_y(0,0,s,t_o(s)); \\ \frac{dt_o(s)}{ds} &= \frac{1}{v_o(s)} = \frac{H_o(0,0,s,t_o(s)) - e\varphi_o(0,0,s,t_o(s))}{p_o(s)c^2} \equiv \frac{E_o(s)}{p_o(s)c^2}; \\ \frac{dE_o(s)}{ds} &\equiv eE_s(0,0,s,t_o(s)). \end{aligned} \quad (20-7)$$

We also introduced longitudinal Canonical pair which has zero values for reference particle:

$$\{\tau = -c(t - t_o(s)), \quad \delta = (H - E_o(s) - e\varphi_o(s, t))/c\} \quad (20-8)$$

Finally, we had derived the complete expression for Hamiltonian, and also the linearized Hamiltonian for an arbitrary accelerator expanding about the point $f(x=0, y=0, s, t_o(s)) \rightarrow f|_{ref}$ in time and space. In normalized Canonical coordinates with $p_{norm} = mc$:

$$\begin{aligned}
 \pi_x &= \frac{P_1}{mc}; \quad \pi_y = \frac{P_3}{mc}; \quad \pi_\tau = \frac{\delta}{mc}; \\
 \mathcal{H}_L &= \frac{mc}{p_o} \cdot \frac{\pi_x^2 + \pi_y^2}{2} + \left(\frac{mc}{p_o} \right)^3 \cdot \frac{\pi_\tau^2}{2} + \\
 &\quad L(x\pi_y - y\pi_x) + g_x x \pi_\tau + g_y y \pi_\tau + \\
 &\quad + \frac{F}{mc} \frac{x^2}{2} + \frac{N}{mc} xy + \frac{G}{mc} \frac{y^2}{2} + \frac{U}{mc} \frac{\tau^2}{2} + \frac{F_x}{mc} x\tau + \frac{F_y}{mc} y\tau;
 \end{aligned} \tag{20-9}$$

where we separated parts of the Hamiltonian into three lines: quadratic form of momenta (“kinetic”), products of momenta and coordinates (mixed) and quadratic form of coordinates (“potential”).

The coefficients of the Hamiltonian are:

$$K = -\frac{e}{p_o c} \left(B_y + \frac{c}{v_o} E_x \right); \quad c_p = \frac{mc}{p_o} = \frac{1}{\beta_o \gamma_o};$$

$$f = \frac{F}{mc} = -\textcolor{blue}{K} \cdot \frac{e}{mc^2} \left(\textcolor{blue}{B}_y + \frac{2c}{v_o} \textcolor{blue}{E}_x \right) - \frac{e}{mc^2} \left(\frac{\partial B_y}{\partial x} + \frac{\partial E_x}{\partial x} \right) + \frac{mc}{p_o} \left(\frac{\textcolor{blue}{e} B_s}{2mc^2} \right)^2 + \frac{\textcolor{red}{mc}}{p_o} \left(\frac{\textcolor{red}{e} E_x}{p_o c} \right)^2;$$

$$g = \frac{G}{mc} = \frac{e}{mc^2} \left(\frac{\partial B_x}{\partial y} - \frac{c}{v_o} \frac{\partial E_y}{\partial y} \right) + \frac{mc}{p_o} \left(\frac{\textcolor{blue}{e} B_s}{2mc^2} \right)^2 + \frac{\textcolor{red}{mc}}{p_o} \left(\frac{\textcolor{red}{e} E_z}{p_o c} \right)^2; \quad ; (20-10)$$

$$2n = \frac{2N}{mc} = \frac{e}{mc^2} \left(\left(\frac{\partial B_x}{\partial x} - \frac{\partial B_y}{\partial y} \right) - \frac{c}{v_o} \left(\frac{\partial E_x}{\partial y} + \frac{\partial E_y}{\partial x} \right) \right) + \textcolor{red}{K} \cdot \frac{e}{mc^2} \left(B_x - \frac{2c}{v_o} E_y \right) + \frac{mc}{p_o} \left(\frac{\textcolor{red}{e} E_z}{p_o c} \right) \left(\frac{\textcolor{red}{m} e E_x}{p_o c} \right);$$

$$L = \kappa + \frac{e}{2p_o c} B_s; \quad u = \frac{U}{mc} = \frac{e}{mc v_o} \frac{\partial E_s}{\partial ct}; \quad g_x = \left(\frac{\textcolor{red}{mc}}{p_o} \right)^2 \frac{\textcolor{red}{e} E_x}{p_o c} - \frac{mc^2}{p_o v_o} K; \quad g_y = \left(\frac{\textcolor{red}{mc}}{p_o} \right)^2 \frac{\textcolor{red}{e} E_y}{p_o c};$$

$$f_x = \frac{F_x}{mc} = \frac{e}{mc^2} \frac{\partial B_y}{\partial ct} + \frac{e}{mc v_o} \frac{\partial E_x}{\partial ct}; \quad f_y = \frac{F_y}{mc} = -\frac{e}{mc^2} \frac{\partial B_x}{\partial ct} + \frac{e}{mc v_o} \frac{\partial E_y}{\partial ct}.$$

Note, that the Hamiltonian (20-9) is dimensionless and its coefficients are either dimensionless or have dimension of 1/L or 1/L². Not all coefficients are important in all case. For example, in ultra-relativistic case, high powers of $mc/p_o = 1/\beta_o \gamma_o$ can be neglected. First will disappear terms in red, then in blue terms could become weak (but not always negligible – beware of this!)

We can easily write matrix form of the Hamiltonian and **D**-matrix and to derive cubic equation for its eigen values. We discussed that eigen values of **D**-matrix are coming in pairs with of eigen values with opposite sign, with reduces equation to a bi-quadratic of power n:

$$p(\lambda) = \prod_{k=1}^3 (\lambda - \lambda_k)(\lambda + \lambda_k) = \prod_{k=1}^3 (\lambda^2 - \lambda_k^2) = \lambda^6 + a_4 \lambda^4 + a_2 \lambda^2 + a_0 = 0 \quad (20-11)$$

In detail:

$$\mathbf{H} = \begin{pmatrix} f & 0 & n & L & f_x & g_x \\ 0 & \frac{mc}{p_o} & -L & 0 & 0 & 0 \\ n & -L & g & 0 & f_y & g_y \\ L & 0 & 0 & \frac{mc}{p_o} & 0 & 0 \\ f_x & 0 & f_y & 0 & u & 0 \\ g_x & 0 & g_y & 0 & 0 & \left(\frac{mc}{p_o}\right)^3 \end{pmatrix}; \quad \mathbf{D} = \begin{pmatrix} 0 & \frac{mc}{p_o} & -L & 0 & 0 & 0 \\ -f & 0 & -n & -L & -f_x & -g_x \\ L & 0 & 0 & \frac{mc}{p_o} & 0 & 0 \\ -n & L & -g & 0 & -f_y & -g_y \\ g_x & 0 & g_y & 0 & 0 & \left(\frac{mc}{p_o}\right)^3 \\ -f_x & 0 & -f_y & 0 & -u & 0 \end{pmatrix}. \quad (20-12)$$

One better use Mathematica to get expressions for coefficients: some are relatively short...

$$a_4 = \sum_{k=1}^3 \lambda_k^2 = 2L^2 + \frac{mc}{p_o} \left(f + g + \left(\frac{mc}{p_o} \right)^2 u \right); c_p \equiv \frac{mc}{p_o};$$

$$a_2 = -(\lambda_1^2 \lambda_2^2 + \lambda_1^2 \lambda_3^2 + \lambda_2^2 \lambda_3^2) = L^4 - c_p^4 (f_x^2 + f_y^2) + 2c_p^3 L^2 u + c_p^2 (f \cdot g - n^2) + c_p (4L(f_x g_y - f_y g_x) - L^2(f + g) + u(f + g - g_x^2 - g_y^2)) \quad (20-13)$$

but the main coefficient is rather long and ugly:

$$a_o = Det[\mathbf{D}] = \prod_{k=1}^3 (-\lambda_k^2) = -\lambda_1^2 \lambda_2^2 \lambda_3^2 =$$

$$c_p^5 (f \cdot g \cdot u + 2nf_x f_y - f \cdot f_y^2 - g \cdot f_x^2 - n^2 u) + c_p^4 L^2 (f_x^2 + f_y^2 + f \cdot u + g \cdot u) \quad (20-14)$$

$$+ c_p^3 L^4 u + c_p^2 \left((f_y g_x - f_x g_y)^2 + u(2g_x g_y n - g \cdot g_x^2 - f \cdot g_y^2) \right) + c_p L^2 u (g_x^2 + g_y^2).$$

Naturally, the cubic equation can be solved analytically:

$$\lambda_1^2 = -\frac{1}{6} \left(2a_4 + \frac{2^{4/3} (3a_2 - a_4^2)}{q} - 2^{2/3} q \right);$$

$$\lambda_{2,3}^2 = -\frac{1}{6} \left(2a_4 - \frac{2^{1/3} (1 \pm i\sqrt{3}) (3a_2 - a_4^2)}{q} \mp i \cdot 2^{-1/3} (i + \sqrt{3}) q \right); \quad (20-15)$$

$$q = \left(3\sqrt{3} \sqrt{27a_0^2 + 4a_2^3 - 18a_0 a_2 a_4 - a_2^2 a_4^2 + 4a_0 a_4^3 - 27a_0 + 9a_2 a_4 - 2a_4^3} \right)^{1/3}.$$

Even though, all of these expressions are explicit and analytical, it still preferable to give a computer to crack the numbers... and to use them to calculate step by step matrices using the most general Sylvester formula. It is especially true for case when we have time-dependent or accelerating fields and can not no longer rely on piece-wise constancy of the Hamiltonian matrix. Meanwhile, computes still can split the steps in sufficiently short steps in s and calculate the transport matrix (using Sylvester formula for exact symplecticity!) to an arbitrary good accuracy. For periodic system (such as a storage ring), one than can solve cubic equation on its eigen values

$$\det[T(s) - \lambda I] = 0 \rightarrow (\lambda + \lambda^{-1})^3 + b_2(\lambda + \lambda^{-1})^2 + b_1(\lambda + \lambda^{-1}) + b_0 = 0 \quad (20-16)$$

$$p_6(\lambda) = \prod_{k=1}^3 (\lambda_k^{-1} - \lambda)(\lambda_k - \lambda) = 0;$$

$$(\lambda_k^{-1} - \lambda) = \left(\frac{\lambda_k}{\lambda} \right) = -(\lambda_k - \lambda^{-1}); (\lambda_k - \lambda^{-1})(\lambda_k - \lambda) = (\lambda_k^2 - \lambda_k(\lambda + \lambda^{-1}) + 1);$$

$$\frac{1}{\lambda^2} p_6(\lambda) = \prod_{k=1}^3 \left((\lambda + \lambda^{-1}) - \frac{\lambda_k^2 - 1}{\lambda_k} \right) = \tilde{p}_3(\lambda + \lambda^{-1}) \rightarrow \tilde{p}_3(\lambda + \lambda^{-1}) = 0$$

$$\lambda^6 + a_5 \lambda^5 + a_4 \lambda^4 + a_3 \lambda^3 + a_2 \lambda^2 + a_1 \lambda + 1 = 0$$

$$\lambda^{-6} + a_5 \lambda^{-5} + a_4 \lambda^{-4} + a_3 \lambda^{-3} + a_2 \lambda^{-2} + a_1 \lambda^{-1} + 1 = 0$$

$$1 + a_5 \lambda + a_4 \lambda^2 + a_3 \lambda^3 + a_2 \lambda^4 + a_1 \lambda^5 + \lambda^6 = 0 \Rightarrow$$

$$a_1 = a_5 = -\text{Trace}[D]; a_2 = a_4.$$

and check that the 3D motion is stable

$$|\lambda_k| = 1; \lambda_k = e^{i\mu_k}; \mu_k = 2\pi Q_k; k = 1, 2, 3 \quad (20-17)$$

and define three eigen vectors and their complex conjugates:

$$Y_k(s) = \begin{bmatrix} w_{kx} e^{i\chi_{kx}} \\ \left(v_{kx} + i \frac{q_{kx}}{w_{kx}} \right) e^{i\chi_{kx}} \\ w_{ky} e^{i\chi_{ky}} \\ \left(v_{ky} + i \frac{q_{ky}}{w_{ky}} \right) e^{i\chi_{ky}} \\ w_{k\tau} e^{i\chi_{k\tau}} \\ \left(v_{k\tau} + i \frac{q_{k\tau}}{w_{k\tau}} \right) e^{i\chi_{k\tau}} \end{bmatrix}; Y_k(s+C) = Y_k(s); T(s)Y_k(s) = e^{i\mu_k} Y_k(s); k = 1, 2, 3 \quad (2-18)$$

with the symplectic orthogonality relations that we already discussed:

$$Y_k^T S Y_j = 0; Y_j^{*T} S Y_k = 2i\delta_{kj}; \quad (2-19)$$

which will apply multiple (15 to be exact!) relations on the component of the eigen vectors, with the simples being:

$$q_{kx} + q_{ky} + q_{k\tau} = 1; k = 1, 2, 3 \quad (2-20)$$

Frequently there are a lot simpler cases, some of which we going to consider.

Accelerator with constant energy – closed orbit.

One of the most used approximations (and simplification) is coming from the fact that in the most of the accelerators (especially in storage ring) longitudinal (or so called synchrotron) oscillations are very slow, when compared with transverse (or so called betatron) oscillations. Specifically, in most of typical storage rings it takes from few hundreds to few thousands of turns to complete one oscillation. Furthermore, in hadron storage ring, where losses on synchrotron are practically absent, one can operate beam in so-called coasting mode – e.g. without any AC fields. Thus, let's consider such an accelerator and study how particles motion depends on their energy (momentum p_o) and explicitly no time dependence.

$$\mathcal{H}_L = \frac{mc}{p_o} \cdot \frac{\pi_x^2 + \pi_y^2}{2} + \left(\frac{mc}{p_o} \right)^3 \cdot \frac{\pi_\tau^2}{2} +$$

$$L(x\pi_y - y\pi_x) + g_x x \pi_\tau + g_y y \pi_\tau + \frac{F}{mc} \frac{x^2}{2} + \frac{N}{mc} xy + \frac{G}{mc} \frac{y^2}{2}; \quad (20-21)$$

$$\pi_x = \frac{P_1}{mc}; \quad \pi_y = \frac{P_3}{mc}; \quad \pi_\tau = \frac{\delta}{mc};$$

Since the energy of the particle is constant but time is slipping:

$$\frac{d}{ds} \pi_\tau = - \frac{\partial H}{\partial \tau} = 0 \rightarrow \pi_\tau = const;$$

$$\frac{d}{ds} \tau_\tau = \frac{\partial H}{\partial \pi_\tau} = g_x x + g_y y + \left(\frac{mc}{p_o} \right)^3 \cdot \pi_\tau \quad (20-22)$$

we can simplify the equations of motion for 2D case plus energy dependence and time slippage:

$$\begin{aligned}\mathcal{H}_L &= \mathcal{H}_\beta + \mathcal{H}_\delta; \quad Z^T = (x, \pi_x, y, \pi_y); \\ \mathcal{H}_\beta &= \frac{mc}{p_o} \cdot \frac{\pi_x^2 + \pi_y^2}{2} + \frac{F}{mc} \frac{x^2}{2} + \frac{N}{mc} xy + \frac{G}{mc} \frac{y^2}{2} + L(x\pi_y - y\pi_x); \\ \mathcal{H}_\delta &= \left(\frac{mc}{p_o} \right)^3 \cdot \frac{\pi_\tau^2}{2} + g_x x \pi_\tau + g_y y \pi_\tau;\end{aligned}\tag{20-23}$$

$$\frac{d}{ds} Z = D_\beta \cdot Z + \pi_\tau \cdot F_\delta; \quad \pi_\tau \cdot C = S \frac{\partial}{\partial Z} H_\delta; \quad F_\delta^T = C \cdot \begin{bmatrix} 0 & -g_x & 0 & -g_y \end{bmatrix};$$

or in explicit matrix form:

$$\begin{aligned}\frac{dZ}{ds} &= D \cdot Z + \pi_\tau \cdot C; \quad D = \begin{bmatrix} 0 & 1 & -L & 0 \\ -f & 0 & -n & -L \\ L & 0 & 0 & 1 \\ -n & L & -g & 0 \end{bmatrix}; \quad C = \begin{bmatrix} 0 \\ -g_x \\ 0 \\ -g_y \end{bmatrix}. \\ \frac{d\tau}{ds} &= g_x x + g_y y + \left(\frac{mc}{p_o} \right)^3 \pi_\tau; \quad g_x = \left(\frac{mc}{p_o} \right)^2 \frac{eE_x}{p_o c} - \frac{mc^2}{p_o v_o} K; \quad g_y = \left(\frac{mc}{p_o} \right)^2 \frac{eE_y}{p_o c}\end{aligned}\tag{20-24}$$

We shall note that for ultra-relativistic particles (or in the absence of the electric fields!) only the curvature K of the trajectory remains as the driving term g_x for transverse motion. Solution for of the in-homogeneous equation for Z can be trivially expressed using 4x4 transport matrix:

$$\begin{aligned}
 Z &= Z_\beta + \pi_\tau \cdot R; \quad \frac{dZ_\beta}{ds} = D \cdot Z_\beta; \quad \frac{dR}{ds} = D \cdot R + C; \\
 \mathbf{M} &\equiv \mathbf{M}_{4 \times 4}; \quad \frac{d\mathbf{M}(s)}{ds} = D \cdot \mathbf{M}(s); \quad Z_\beta(s) = \mathbf{M}(s) Z_{\beta o}; \\
 R(s) &= \mathbf{M}(s) A(s); \quad \mathbf{M}(s) \frac{dA}{ds} = \pi_\tau \cdot C \Rightarrow \frac{dA}{ds} = \mathbf{M}^{-1}(s) C(s); \mathbf{M}^{-1} = -\mathbf{S} \mathbf{M}^T \mathbf{S}; \\
 \Rightarrow A(s) &= \int_o^s \mathbf{M}^{-1}(\xi) C(\xi) d\xi; \quad R(s) = \mathbf{M}(s) \left(A_o + \int_o^s \mathbf{M}^{-1}(\xi) C(\xi) d\xi \right) \\
 R(s) &= \int_o^s \mathbf{M}(\xi|s) C(\xi) d\xi; \eta
 \end{aligned} \tag{20-25}$$

For periodic system we can find “periodic transverse orbit” for an off-momentum particle:

$$\begin{aligned}
 \eta(s+C) &= \int_o^{s+C} \mathbf{M}(\xi|s+C)C(\xi)d\xi = \mathbf{T}(s)\eta(s) + \int_s^{s+C} \mathbf{M}(\xi|s+C)C(\xi)d\xi \\
 \mathbf{M}(\xi|s+C) &= \mathbf{T}(s)\mathbf{M}(\xi|s); \mathbf{T}(s) \equiv \mathbf{M}(s|s+C); \\
 \eta(s+C) = \eta(s) &\Rightarrow (\mathbf{I} - \mathbf{T})\eta(s) = \int_s^{s+C} \mathbf{M}(\xi|s+C)C(\xi)d\xi; \\
 \eta(s) &= (\mathbf{I} - \mathbf{T}(s))^{-1} \int_s^{s+C} \mathbf{M}(\xi|s+C)C(\xi)d\xi; \quad Z = Z_\beta + \pi_\tau \cdot \eta(s). \quad .
 \end{aligned} \tag{20-26}$$

We already found expression for such closed periodical orbit expressed via eigen vectors – naturally the results are identical. The $\eta = \begin{bmatrix} \eta_x & \eta_{px} & \eta_y & \eta_{py} \end{bmatrix}$ - function is called transverse dispersion (picking analogy from optics). Unfortunately in accelerator physics terminology there is a number of confusions... and frequently the dispersion is represented by $D = \begin{bmatrix} D_x & D_{px} & D_y & D_{py} \end{bmatrix}$. Read the context to be sure...

Next natural step is to look onto the slippage of the particle in time for a particle without betatron oscillations $Z_\beta=0$ (we will add them later):

$$Z = \pi_\tau \cdot \eta(s); \frac{d\tau}{ds} = \left(g_x \eta_x + g_y \eta_y + \left(\frac{mc}{p_o} \right)^3 \right) \pi_\tau; \quad (20-27)$$

$$\tau(s) = f_\tau(s) \pi_\tau; \quad f_\tau(s) = f_\tau(0) + \left(\frac{mc}{p_o} \right)^3 \cdot s + \int_0^s (g_x(\xi) \eta_x(\xi) + g_y(\xi) \eta_y(\xi)) d\xi.$$

First (red) term corresponds to the velocity dependence on the particles energy – it is weak for ultra-relativistic particles moving very-very close to the speed of the light, but it is important for hadron accelerators. Hence, we will keep it. Again, for a periodic system we

$$f_\tau(s+C) = f_\tau(s) + \left(\frac{mc}{p_o} \right)^3 \cdot C + \int_s^{s+C} (g_x(\xi) \eta_x(\xi) + g_y(\xi) \eta_y(\xi)) d\xi = \eta_\tau \cdot C; \quad (20-28)$$

$$\eta_\tau = \frac{1}{C} \int_0^C (g_x \eta_x + g_y \eta_y) ds + \left(\frac{mc}{p_o} \right)^3;$$

It worth expressing it for a simple case when electric field is zero

$$\eta_\delta = \frac{p_o}{mc} \eta_\tau = \left(\frac{mc}{p_o} \right)^2 - \frac{1}{C} \frac{c}{v_o} \int_0^C K(s) \eta_x(s) ds = \frac{1}{\beta_o^2 \gamma_o^2} - \frac{1}{\beta_o} \langle K \eta_x \rangle \quad (20-28)$$

e.g. the dependence of the travel time around the storage ring on particles momentum as two components: one corresponds to increase in velocity (kinematic) and the other (geometrical) to - typically - elongation of the trajectory in bending magnets – particles with higher energy travel at larger radius. In general, η_τ can be either negative or positive. When two terms cancel each other, travel time around storage ring is energy independent – this energy is called critical. If the geometrical term $\langle g_x \eta_x + g_y \eta_y \rangle$ in the accelerator is positive, the accelerator does not have critical energy. Such conditions do not come naturally and require a special, so call negative compaction factor lattice:

$$\alpha_c = \langle g_x \eta_x + g_y \eta_y \rangle > 0 .$$

Synchrotron oscillations.

Here we will assume that longitudinal oscillations (if stable) are slow. Let's initially introduce longitudinal field

$$u = -\frac{e}{p_o c^2} \frac{\partial E_{rf}}{\partial t}$$

and see how it affects betatron oscillations. Assuming that we can use perturbation theory we have

$$\begin{aligned} \delta\phi_k &= \frac{1}{2} \int_0^s \tilde{Y}_k^{*T}(\zeta) \Delta \mathbf{H}(\zeta) \tilde{Y}_k(\zeta) d\zeta; \\ \Delta \mathbf{H}_{55} &= -\frac{e}{p_o c^2} \frac{\partial E_{rf}}{\partial t}; \delta Q_k = -\frac{1}{4\pi} \frac{e}{p_o c} \int_0^c \frac{\partial E_{rf}}{\partial ct} |Y_{k5}|^2 ds \end{aligned} \quad (20-29)$$